## Introduction to Quantum Mechanic

A) Radiation
B) Light is made of particles. The need for a quantification

1) Black-body radiation (1860-1901)
2) Atomic Spectroscopy (1888-)
3) Photoelectric Effect (1887-1905)
C) Wave-particle duality
4) Compton Effect (1923).
5) Electron Diffraction Davisson and Germer (1925).
6) Young's Double Slit Experiment
D) Louis de Broglie relation for a photon from relativity
E) A new mathematical tool: Wavefunctions and operators
F) Measurable physical quantities and associated operators -

Correspondence principle
G) The Schrödinger Equation (1926)
H) The Uncertainty principle

When you find this image,
you may skip this part
This is less important


Christiaan Huygens Dutch 1629-1695 light consists of waves

The idea of duality is rooted in a debate over the nature of light and matter dating back to the 1600s, when competing theories of light were proposed by Huygens and Newton.

## Sir Isaac Newton

 16431727light consists of particles

## Radiations, terminology

Considering radiations as waves: a periodic function $\Psi(r, t)$ where $r$ is and $t$ is time. Instead of cosine and sine we will use equivalent exponential expressions:

$$
\begin{array}{ll}
\mathrm{e}^{i x}=\cos x+i \sin x & \rightarrow \\
\cos x=\frac{e^{i x}+e^{-i x}}{2} \\
e^{-i x}=\cos x-i \sin x & \rightarrow \\
\sin x=\frac{e^{i x}-e^{-i x}}{2 i}
\end{array}
$$

Two expressions $\left.\quad \Psi=A e^{i(k r-\omega t)}=A e^{2 \pi i\left({ }^{\perp}\right.}-v t\right)$
A is the amplitude
The beam intensity is given by $\Psi^{*} \Psi=A^{2}$
which depends neither from k , nor from $\lambda, \nu$, and $\omega$.
$\boldsymbol{\lambda}$ is the wave length (dimension of a length);
$\mathrm{k}=\frac{2 \pi}{\lambda}$ is the wave number (inverse of a length)
$v$ is the frequency; $\omega=2 \pi v$ is angular frequency (inverse of time).

## Interferences

$\Psi$ is periodic; $e^{i 2 \pi}=1$
Adding $2 \pi$ to the exponent (either by increasing $r$, or $t$ ), the wave remains unaffected. Two waves are in phase for $\mathrm{t}=0$

- if $\frac{\mathbf{r}_{2}}{\lambda_{2}}-\frac{\mathbf{r}_{1}}{\lambda_{1}}=\mathbf{n}$

$$
\text { Or if } \quad \mathrm{k}_{2} \mathrm{r}_{2}-\mathrm{k}_{1} \mathrm{r}_{1}=2 \pi \mathrm{n}
$$

Two waves are in phase at the origin $(\mathrm{r}=0$ )

- $\quad v_{2} t_{2}-v_{1} t_{1}=n$

$$
\begin{aligned}
& \text { if } \frac{t_{2}}{T_{2}}-\frac{t_{1}}{T_{1}}=n \\
& \text { or if } \omega_{2} t_{2}-\omega_{1} t_{1}=2 \pi n
\end{aligned}
$$

with $n$ integer

Constructive Interferences


Destructive Interferences


## Phase speed or velocity

$$
\left(\mathrm{kr}_{2}-\omega \mathrm{t}_{2}\right)=\left(\mathrm{kr}_{1}-\omega \mathrm{t}_{1}\right) \quad \rightarrow \quad \mathrm{k}\left(\mathrm{r}_{2}-\mathrm{r}_{1}\right)=\omega\left(\mathrm{t}_{2}-\mathrm{t}_{1}\right)
$$

Where from $\mathrm{v}_{\varphi}=\frac{\mathrm{r}_{2}-\mathrm{r} 1}{\mathrm{t}_{2}-\mathrm{t}_{1}}=\frac{\omega}{\mathrm{k}}=\lambda \nu \quad$ et $v=\frac{\mathrm{v}_{\varphi}}{\lambda}$
Warning, as we will see later

$$
\begin{aligned}
& \frac{\Delta \mathrm{r}}{\Delta \mathrm{t}} \text { is not equal to } \frac{\mathrm{dr}}{\mathrm{dt}} \text { ! } \\
& \mathrm{v}_{\varphi} \text { is not equal to } \mathrm{v}=\frac{\mathrm{dr}}{\mathrm{dt}}
\end{aligned}
$$

The velocity, v , will be defined as a derivative :
$\mathrm{v}=\frac{\mathrm{dr}}{\mathrm{dt}}=\frac{\mathrm{d} \omega}{\mathrm{dk}} . \quad \mathrm{v}_{\varphi}=\frac{\Delta \mathrm{r}}{\Delta \mathrm{t}} \quad$ would be equal to $\mathrm{v}=\frac{\mathrm{dr}}{\mathrm{dt}}=\frac{\omega}{\mathrm{k}} \quad$ only if k vs. $\omega$ is a linear expression which is not generally true. ।

$$
\text { For a photon } \mathrm{v}_{\varphi}=\mathrm{c}=\lambda v \text { et } v=\frac{\mathrm{c}}{\lambda}
$$

## Introducing new variables

- At the moment, let consider this just a formal change, introducing

$$
\mathrm{p}=\frac{\mathrm{h}}{\lambda} \text { and } \mathrm{E}=\mathrm{h} v
$$

and

$$
\begin{aligned}
& \mathrm{K}=\frac{\mathrm{h}}{2 \pi} \quad \text { we obtain } \\
& \Psi=A \mathrm{e}^{\mathrm{i}(\mathrm{kr}-\omega \mathrm{t})}=\mathrm{A} \mathrm{e}^{2 \pi \mathrm{i}\left(\frac{\mathrm{r}}{\lambda}-v \mathrm{t}\right)}=\mathrm{Ae} \mathrm{i}_{(\mathrm{hr}-\mathrm{Et})}^{\lambda}
\end{aligned}
$$

## Introducing new variables

$$
\left.\Psi=A \mathrm{e}^{\mathrm{i}(\mathrm{kr}-\omega \mathrm{t})}=\mathrm{A} \mathrm{e}^{2 \pi \mathrm{i}\left({ }^{\mathrm{r}}\right.}-\mathrm{Vt}\right)=\mathrm{Ae} \mathrm{e}_{\mathrm{h}}^{\mathrm{i}}(\mathrm{pr}-\mathrm{Et})
$$

At the moment, h is a simple constant Later on, $h$ will have a dimension and the $p$ and $E$ will be physical quantities
Then

$$
\mathrm{p}=\mathrm{hk}=\frac{\mathrm{h}}{\lambda} \quad ; \quad \mathrm{E}=\mathrm{h} \omega=\mathrm{h} \nu=\frac{\mathrm{h}}{\mathrm{~T}} \quad \text { et } \mathrm{v}_{\varphi}=\frac{\mathrm{E}}{\mathrm{p}}
$$

## 2 different velocities, v and v

$$
v_{\varphi}=\frac{E}{p}
$$

$$
\mathrm{E}=\frac{\mathrm{mv}}{}{ }^{2} \text { and } \mathrm{p}=\mathrm{mv} \rightarrow \quad \mathrm{v}=\frac{2 \mathrm{E}}{\mathrm{p}}=2 \mathrm{v} \varphi
$$

${ }^{\mathrm{v}}$ differs from $\mathrm{v}_{\varphi}$.

## If h is the Planck constant J.s

## Then



## electron discovery in 1887 <br> - J. J. Thomson (1887) Some negative part could be extracted from the atoms

- Robert Millikan (1910) showed that it was quantified.
-Rutherford (1911) showed that the negative part was diffuse while the positive part was concentrated.



## black-body radiation

Gustav Kirchhoff (1860). The light emitted by a black body is called black-body radiation

At room temperature, black bodies emit $\operatorname{IR}$ light, but as the temperature increases past a few hundred degrees Celsius, black bodies start to emit at visible wavelengths, from red, through orange, yellow, and white before ending up at blue, beyond which the emission includes increasing amounts of UV


## black-body radiation

こlassical Theory
Fragmentation of the surface.
One large area (Small $\lambda$ Large $\mathbf{v}$ ) $\rightarrow$ smaller pieces (Large $\lambda$ Small $\mathbf{v}$ )
Vibrations associated to the size, $\mathrm{N}^{2}$ or $\mathrm{N}^{3}$


## black-body radiation

## Kirchhoff

Radiation is emitted when a solid after receiving energy goes back to the most stable state (ground state). The energy associated with the radiation is the difference in energy between these 2 states. When T increases, the average $\mathbf{E}_{\text {Mean }}^{*}$ is higher and intensity increases.
$\mathrm{E}^{*}{ }_{\text {mean }}-\mathrm{E}=\mathrm{kT}$.
k is Boltzmann constant ( $\mathrm{k}=1.38$ $10^{-23}$ Joules $\mathrm{K}^{-1}$ ).


RED
Small v

Why a decrease for small $\lambda$ ?
Quantification

## black-body radiation



Max Planck (1901) Göttingen

Numbering rungs of ladder introduces quantum numbers (hete equally spaced)

## Quantum numbers

In mathematics, a natural number (also called counting number) has two main purposes: they can be used for counting ("there are 6 apples on the table"), and they can be used for ordering ("this is the 3rd largest city in the country").


Why a decrease for small $\lambda$ ?

## Quantification

## black-body radiation



Max Planck (1901) Göttingen



## black-body radiation, quantification

Max Planck



Steps too hard to climb Pyramid nowadays
 classical theory ( 5000 K )


Easy slope, ramp
Pyramid undergconstruction

Everything depends on the ratio $\frac{U}{T}$

$$
E_{\text {average }}^{*}-E=\frac{h \nu}{\exp \left(\frac{\mathrm{~h} v}{\mathrm{kT}}\right)-1}
$$

If $\frac{v}{T}$ is weak $e^{x}=1+x+\ldots$

$$
\mathbf{E}_{\text {average }}-\mathbf{E}=\frac{\mathrm{h} \nu}{\mathrm{e}^{\mathrm{X}}-1}=\frac{\mathrm{h} \nu}{\mathrm{x}}=\mathrm{kT}
$$

$\mathrm{E}^{*}{ }_{\text {average }}-\mathrm{E}$ is that given by classical theory.
If $\frac{U}{T}$ is high (high frequency - low T),
$E^{*}$ average $-E$ tends to 0 . The radiation intensity tends to 0 .


Atomic Spectroscopy

## Absorption or Emission

$$
\frac{1}{\lambda_{\mathrm{vac}}}=R_{\mathrm{H}}\left(\frac{1}{n_{1}^{2}}-\frac{1}{n_{2}^{2}}\right)
$$

Johannes Rydberg 1888 Swedish

| $n_{1} \rightarrow n_{2}$ | name | Converges <br> to (nm) |
| :---: | :---: | :---: |
| $1 \rightarrow \infty$ | Lyman | 91 |
| $2 \rightarrow \infty$ | Balmer | 365 |
| $3 \rightarrow \infty$ | Pashen | 821 |
| $4 \rightarrow \infty$ | Brackett | 1459 |
| $5 \rightarrow \infty$ | Pfund | 2280 |
| $6 \rightarrow \infty$ | Humphreys | 3283 |



Atomic Spectroscopy

## Absorption or Emission

Johannes Rydberg 1888 Swedish

$$
\frac{1}{\lambda_{\mathrm{vac}}}=R_{\mathrm{H}}\left(\frac{1}{n_{1}^{2}}-\frac{1}{n_{2}^{2}}\right)
$$

## Emission

Quantum numbers n, levels are not equally spaced
$21 \mathrm{R}=13.6 \mathrm{eV}$

## Photoelectric Effect (1887-1905)

discovered by Hertz in 1887 and explained in 1905 by Einstein.


Heinrich HERTZ (1857-1894)


Albert EINSTEIN (1879-1955)


Every individual photon interacts with the metal surface. This can only be effective if it has the necessary energy $\mathrm{Emin}_{\text {m }}$ to wrest the electron of the metal less strongly hold that has quantified energy level, $E_{\text {min }}$. The frequency threshold is therefore $v_{0}=\frac{\mathrm{E}_{\min }}{h}$. You can not combine the energy from two photons to remove electrons: below $v \gg v_{0}$ intensity is zero. If the radiation has a frequency $v \gg v_{0,}$, the kinetic energy of the electron ripped off is the excess energy: $\mathrm{E}_{\mathrm{kin}}=\mathrm{h} \nu-\mathrm{E}_{\mathrm{min}}$. This energy is proportional to $v$.

## Compton effect 1923 <br> playing billiards assuming $\lambda=h / p$



Arthur Holly Compton American
1892-1962

$$
\left(v=\frac{c}{\lambda}\right) \quad \mathrm{p}
$$

Energy Conservation
$\mathrm{h} \frac{c}{\lambda}=h \frac{c}{\lambda^{\prime}}+\frac{p^{2}}{2 m}$
Momentum Conservation (projection on $x$ )
$\frac{\mathbf{h}}{\lambda}=\frac{\mathbf{h}}{\lambda^{\prime}} \cos \theta+\mathrm{p} \cos \alpha$
Momentum Conservation (projection on y)
$0=\frac{\mathrm{h}}{\lambda^{\prime}} \sin \theta-\mathrm{p} \sin \alpha$

## Davisson and Germer 1925



Clinton Davisson
$2 \mathrm{~d} \sin \quad \theta=\mathrm{k} \quad \lambda$ Lester Germer In 1927

Diffraction is similarly observed using a mono-energetic electron beam
Bragg law is verified assuming $\lambda=h / p$

## Wave-particle Equivalence.

-Compton Effect (1923).
-Electron Diffraction Davisson and Germer (1925)
-Young's Double Slit Experiment

## Wave-particle duality

In physics and chemistry, wave-particle duality is the concept that all matter and energy exhibits both wave-like and particle-like properties. A central concept of quantum mechanics, duality, addresses the inadequacy of classical concepts like "particle" and "wave" in fully describing the behavior of small-scale objects. Various interpretations of quantum mechanics attempt to explain this apparent paradox.

## Thomas Young 1773-1829



English, was born into a family of Quakers.
At age 2, he could read.
At 7, he learned Latin, Greek and maths.
At 12, he spoke Hebrew, Persian and could handle optical instruments.
At 14, he spoke Arabic, French, Italian and Spanish, and soon the Chaldean Syriac. "...
He is a PhD to 20 years "gentleman, accomplished flute player and minstrel (troubadour). He is reported dancing above a rope."
He worked for an insurance company, continuing research into the structure of the retina, astigmatism ...
He is the rival Champollion to decipher hieroglyphics.
He is the first to read the names of Ptolemy and Cleopatra which led him to propose a first alphabet of hieroglyphic scriptures (12 chảracters).

## Young's Double Slit Experiment



## Young's Double Slit Experiment

This is a typical experiment showing the wave nature of light and interferences.
What happens when we decrease the light intensity ?
If radiation = particles, individual photons reach one spot and there will be no interferences If radiation $\neq$ particles there will be no spots on the screen

The result is ambiguous
There are spots
The superposition of all the impacts make interferences

## Young's Double Slit Experiment

Assuming a single electron each time What means interference with itself ?

What is its trajectory?
If it goes through F1, it should ignore the presence of F2


## Young's Double Slit Experiment

There is no possibility of knowing through which split the photon went! If we measure the crossing through F1, we have to place a screen behind.

Then it does not go to the final screen.
We know that it goes through F1 but we do not know where it would go after.
These two questions are not compatible

Two important differences with classical physics:

- measurement is not independent from observer
- trajectories are not defined; hv goes through F1 and F2 both! or through them with equal probabilities!


Macroscopic world:
A basket of cherries
Many of them (identical)
We can see them and taste others
Taking one has negligible effect Cherries are both red and good

Microscopic world:
A single cherry
Either we look at it without eating
It is red
Or we eat it, it is good
You can not try both at the same time The cherry could not be good and red at the same time


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Slot machine "one-arm bandit"
After introducing a coin, you have 0 coin or $X$ coins.

A measure of the profit has been made: profit = X

## de Broglie relation from relativity

Popular expressions of relativity: $\mathrm{m}_{0}$ is the mass at rest, m in motion

$$
\mathrm{m}=\frac{\mathbf{m}_{0}}{\Gamma} \quad \text { and } \quad E=\mathrm{m} \mathrm{c}^{2}
$$

$E$ like to express $E(m)$ as $E(p)$ with $p=m v$

$$
E^{2}=m^{2} c^{4}\left(1-\frac{v^{2}}{c^{2}}\right)+p^{2} c^{2}=m_{0}^{2} c^{4}+p^{2} c^{2}
$$

$$
\begin{aligned}
&=\sqrt{m_{0}^{2} c^{4}+p^{2} c^{2}}=m_{0} c^{2}\left(1+\frac{p^{2}}{m_{0}^{2} c^{2}}\right)^{1 / 2}= E_{i}+\frac{p^{2}}{2 m_{0}}+\ldots \\
& E_{i}+T+E_{\text {relativistic }}+\ldots \\
& 34
\end{aligned}
$$

## de Broglie relation from relativity



$$
E^{2}=m_{0}^{2} c^{4}+p^{2} c^{2}
$$

Application to a photon $\left(\mathrm{m}_{0}=0\right)$


To remember


## Useful to remember to relate energy and wavelength

## Max Planck

$$
\lambda=\frac{\text { hc }}{\mathrm{E}} \rightarrow \lambda(\boldsymbol{\Lambda})=\frac{12410}{\mathbf{E}(\mathbf{e} \mathbf{V})}
$$

A New mathematical tool:

## Wave functions and Operators

Each particle may be described by a wave function $\Psi(x, y, z, t)$, real or complex, having a single value when position ( $x, y, z$ ) and time ( t ) are defined.
If it is not time-dependent, it is called stationary.
The expression $\Psi=A e^{(p r-E t)}$ does not represent one molecule but a flow of particles: a plane wave

## Wave functions describing one particle

To represent a single particle $\boldsymbol{\Psi}(x, y, z)$ that does not evolve in time, $\boldsymbol{\Psi}(x, y, z)$ must be finite ( 0 at $\infty$ ).

In QM, a particle is not localized but has a probability to be in a given volume:
$\mathrm{dP}=\boldsymbol{\Psi}^{*} \boldsymbol{\Psi} \mathrm{dV}$ is the probability of finding the particle in the volume dV . Around one point in space, the density of probability is $\mathrm{dP} / \mathrm{dV}=\boldsymbol{\Psi}^{*} \boldsymbol{\Psi}$ $\Psi$ has the dimension of $L^{-1 / 3}$
Integration in the whole space should give one $\Psi$ is said to be normalized.


## Operators associated to physical quantities

We cannot use functions (otherwise we would end with classical mechanics)
Any physical quantity is associated with an operator. An operator $O$ is "the recipe to transform $\boldsymbol{\Psi}$ into $\boldsymbol{\Psi}$ " "
We write: $\quad O \Psi=\Psi^{\prime}$
If $O \boldsymbol{\Psi}=\mathrm{o} \boldsymbol{\Psi}$ ( o is a number, meaning that O does not modify $\boldsymbol{\Psi}$, just a scaling factor), we say that $\Psi$ is an eigenfunction of $O$ and $o$ is the eigenvalue. We have solved the wave equation $\mathrm{O} \boldsymbol{\Psi}=o \boldsymbol{\Psi}$ by finding simultaneously $\boldsymbol{\Psi}$ and o that satisfy the equation.
o is the measure of $O$ for the particle in the state described by $\Psi$.


O is a Vending machine (cans) Introducing a coin, you get one can.

No measure of the gain is made unless you sell the can (return to coins)


Slot machine (one-arm bandit)
Introducing a coin, you have 0 coin or $X$ coins.

A measure of the profit has been made: profit $=\mathbf{X}$

## Examples of operators in mathematics: P parity

$$
\operatorname{Pf}(x)=f(-x)
$$

Even function : no change after $x \rightarrow-x$
Odd function : f changes sign after $\mathrm{x} \rightarrow-\mathrm{x}$

$$
\begin{gathered}
y=x^{2} \text { is even } \\
y=x^{3} \text { is odd } \\
y=x^{2}+x^{3} \text { has no parity: } P\left(x^{2}+x^{3}\right)=x^{2}-x^{3}
\end{gathered}
$$

## Examples of operators in mathematics: A

$$
\mathrm{A}=\frac{\mathrm{d}^{2}}{\mathrm{dx}^{2}}-\mathrm{x}^{2}
$$

Apply to the wave function $y=e^{-x^{2} / 2}$

$$
\begin{aligned}
& A=\frac{d^{2} y}{d x^{2}}-x^{2} y=\frac{d^{2} e^{-x^{2} / 2}}{d x^{2}}+x^{2} e^{-x^{2} / 2} \\
& A y=\frac{d\left(-x e^{-x^{2} / 2}\right)}{d x}-x^{2} e^{-x^{2} / 2}=-e^{-x^{2} / 2}+x^{2} e^{-x^{2} / 2}-x^{2} e^{-x^{2} / 2} \\
& A y=-e^{-x^{2} / 2}=-y
\end{aligned}
$$

$y$ is an eigenvector; the eigenvalue is -1

## Linearity

The operators are linear:
$\mathrm{O}\left(\mathrm{a} \Psi_{1}+\mathrm{b} \Psi_{1}\right)=\mathrm{O}\left(\mathrm{a} \Psi_{1}\right)+\mathrm{O}\left(\mathrm{b} \Psi_{1}\right)$

## Normalization

An eigenfunction remains an eigenfunction when multiplied by a constant
$O(\lambda \Psi)=O(\lambda \Psi)$ thus it is always possible to normalize a finite function

$$
\int_{\text {total }} \Psi_{- \text {space }} \Psi^{*} \Psi d V=N \text { taking } \Psi^{\prime}=\frac{1}{\sqrt{N}} \Psi \text { gives } \int_{\text {total_space }} \Psi^{\prime^{*}} \Psi^{\prime} d V=1
$$

Dirac notations < $\Psi \mid \Psi>$

## Mean value

- If $\Psi_{1}$ and $\Psi_{2}$ are associated with the same eigenvalue $\mathrm{o}: \mathrm{O}\left(\mathrm{a} \Psi_{1}+\mathrm{b} \Psi_{2}\right)=\mathrm{o}\left(\mathrm{a} \Psi_{1}+\mathrm{b} \Psi_{2}\right)$
- If not $\mathrm{O}\left(\mathrm{a} \Psi_{1}+\mathrm{b} \Psi_{2}\right)=\mathrm{o}_{1}\left(\mathrm{a} \Psi_{1}\right)+\mathrm{o}_{2}\left(\mathrm{~b} \Psi_{2}\right)$
we define $\bar{o}=\left(a^{2} \mathrm{o}_{1}+b^{2} \mathrm{o}_{2}\right) /\left(\mathrm{a}^{2}+\mathrm{b}^{2}\right)$

$$
\begin{aligned}
& \Psi^{*} \mathbf{O} \Psi=\Psi * \text { o } \Psi \\
& \mathrm{o} \text { is a multiplying factor } \\
& \int_{\infty}\left(\Psi^{*} O \Psi\right) d V=\mathrm{o} \int_{\infty}(\Psi * \Psi) d V \\
& \mathrm{o}_{=}=\frac{\int_{\infty}(\Psi * O \Psi) d V}{\int_{\infty}(\Psi * \Psi) d V}=\frac{\langle\Psi| O|\Psi\rangle}{\langle\Psi \mid \Psi\rangle} \text { Dirac notations }
\end{aligned}
$$

## Sum, product and commutation of

 $(A+B) \Psi=A \Psi+B \Psi$ operators$$
(A B) \Psi=A(B \Psi)
$$

wavefunctions

operators |  | $y_{1}=e^{4 x^{*}}$ | $y_{2}=x^{2}$ | $y_{3}=1 / x$ |
| :---: | :---: | :---: | :---: |
| $d / d x$ | 4 | -- | -- |
| $x 3$ | 3 | 3 | 3 |
| $x d / d x$ | -- | 2 | -1 |

## Sum, product and commutation of operators

$[A, C]=A C-C A \neq 0$
$\begin{aligned} & {[A, B]=A B-B A=0} \\ & {[B, C]=B C-C B=0}\end{aligned} \quad[A, C](y)=A C(y)-C A(y)=\frac{d}{d x}\left[x \frac{d y}{d x}\right]-x\left[\frac{d}{d x}\left(\frac{d y}{d x}\right)\right]$

$$
[A, C](y)=x \frac{d^{2} y}{d x^{2}}+\frac{d y}{d x}-x \frac{d^{2} y}{d x^{2}}=A(y)
$$

$$
[\mathrm{A}, \mathrm{C}]=\mathrm{A}
$$



## Compatibility, incompatibility of operators

$$
\begin{aligned}
& {[A, C]=A C-C A \neq 0} \\
& {[A, B]=A B-B A=0} \\
& {[B, C]=B C-C B=0}
\end{aligned}
$$

When operators commute, the physical quantities may be simultaneously defined (compatibility)
When operators do not commute, the physical quantities can not be simultaneously defined (incompatibility)


## $x$ and $d / d x$ do not commute, are incompatible

$$
\begin{aligned}
& {[d / d x, x]=d / d x(x)-x(d / d x)} \\
& {\left[d / d x, x \rrbracket(y)=\frac{d(x y)}{d x}-x \frac{d y}{d x}=x \frac{d y}{d x}+y-x \frac{d y}{d x}=y\right.} \\
& {[d / d x, x]=+1}
\end{aligned}
$$

Translation and inversion do not commute, are incompatible Translation vector


## Introducing new variables

Now it is time to give a physical meaning. p is the momentum, E is the Energy $\mathrm{H}=6.62 \quad 10^{-34} \mathrm{~J} . \mathrm{s}$

$$
\begin{aligned}
& \mathrm{p}=\frac{\mathrm{h}}{\lambda} \text { and } \mathrm{E}=\mathrm{h} \nu \\
& \mathrm{~h}=\frac{\mathrm{h}}{2 \pi} \\
& \Psi=\mathrm{A} \mathrm{e}^{\mathrm{i}(\mathrm{kr}-\omega \mathrm{t})}=\mathrm{A} \mathrm{e}^{2 \pi \mathrm{i}\left(\frac{\mathrm{r}}{\lambda}-v \mathrm{t}\right)}=\mathrm{A} \mathrm{e}^{\mathrm{i}}(\mathrm{pr}-\mathrm{Et})
\end{aligned}
$$

$$
\Psi=A \mathrm{e}^{\mathrm{i}(\mathrm{kr}-\omega \mathrm{t})}=\mathrm{A} \mathrm{e}^{2 \pi \mathrm{i}\left(\frac{\mathrm{r}}{\lambda}-\nu \mathrm{t}\right)}=\mathrm{A} \mathrm{e}_{\mathrm{K}}^{\mathrm{i}}(\mathrm{pr}-\mathrm{Et})
$$

This represents a (monochromatic) beam, a continuous flow of particles with the same velocity (monokinetic).
$\mathrm{k}, \lambda, \omega, \nu, \mathrm{p}$ and E are perfectly defined R (position) and t (time) are not defined. $\Psi \Psi^{*}=A^{2}=$ constant everywhere; there is no localization.
If $E=$ constant, this is a stationary state, independent of $t$ which is not defined.

## Correspondence principle 1913/1920



> For every physical quantity one can define an operator. The definition uses formulae from classical physics replacing quantities involved by the corresponding operators

Niels Henrik David Bohr Danish
1885-1962

QM is then built from classical physics in spite of demonstrating its limits

## Operators p and H

We use the expression of the plane wave which allows defining exactly $p$ and $E$.

## Momentum and Energy Operators

$$
\begin{aligned}
& \Psi(\mathrm{r}, \mathrm{t})=\Psi(\mathrm{r}) \Psi(\mathrm{t})=\mathrm{Ae}_{\neq}^{\mathrm{i}_{(\mathrm{pr})}} \mathrm{e}_{h}^{\mathrm{i}_{(-\mathrm{Et})}} . \\
& \Psi(\mathrm{r})=\mathrm{Ae}_{h}^{\mathrm{i}_{(\mathrm{pr})}} \quad \Psi(\mathrm{t})=\mathrm{e}_{h}^{\mathrm{i}_{(-\mathrm{Et})}} \\
& \frac{\partial \Psi}{\partial r}=\frac{i p}{h} \Psi \text { and } \frac{\partial \Psi}{\partial t}=-\frac{i \mathrm{E}}{h} \Psi
\end{aligned}
$$

Remember during this chapter

## Stationary state E=constant



Remember for 3 slides after

## Kinetic energy

Classical

$$
\mathrm{T}=\frac{\mathrm{p}^{2}}{2 \mathrm{~m}} \quad \text { quantum operator } \quad \mathbb{P}^{2}=\frac{-\mathbf{h}^{2}}{\mathbf{2 m}} \frac{\overparen{D}^{2}}{\partial \mathrm{I}^{2}}
$$

In 3D :

$$
\mathbb{T}=\frac{\mathbf{- h ^ { 2 }}}{\mathbf{2 m}}\left(\frac{\partial^{2}}{\partial \mathbb{I}^{2}}+\frac{\partial^{2}}{\partial y^{2}}+\frac{\frac{\partial^{2}}{\partial z^{2}}}{2 z^{2}}\right)=\frac{\mathbf{- h ^ { 2 }}}{\mathbf{2 m}} \Delta
$$

Calling

$$
\Delta=\left(\frac{\partial^{2}}{\partial z^{2}}+\frac{\partial^{2}}{\partial y^{2}}+\frac{\theta^{2}}{\partial z^{2}}\right) \quad \text { the laplacian }
$$



Pierre Simon, Marquis de Laplace (1749-1827)

# Correspondence principle angular momentum 

Classical expression
Quantum expression

$$
I_{z}=x p_{y}-y p_{x}
$$

$$
\mathbb{U}_{\mathbb{Z}}=-\mathbf{i h}\left(\mathbb{J} \frac{\partial}{\partial \nabla}=8 \frac{\partial}{\theta \pi}\right)
$$

## Quantum-mechanical Operators

| Operator name | Symbol | Form | Notes |
| :---: | :---: | :---: | :---: |
| $\begin{gathered} \text { position } \\ \text { (in } \mathrm{x} \text { direction, for example) } \end{gathered}$ | $\hat{\mathrm{x}}$ | x | There are also corresponding operators for $y, z$. |
| position (three-dimensional, <br> Cartesian coordinates) | $\hat{\mathbf{r}}$ | $(\mathrm{x}, \mathrm{y}, \mathrm{z})=\mathrm{x} \overrightarrow{\mathrm{x}}+\mathrm{y} \overrightarrow{\mathrm{y}}+\mathrm{z} \overrightarrow{\mathrm{z}}$ |  |
| dipole moment (threedimensional, Cartesian coordinates) | $\hat{\boldsymbol{\mu}}$ | $\sum_{\mathrm{j}=\text { alidams }} q_{j} \mathbf{r}_{\mathrm{j}}$ | q is the charge on each atom; $r$ is the position of each atom. |
| dipole moment (one atom in spherical coordinates) | $\hat{\boldsymbol{\mu}}$ | $-\mathrm{er}(\underline{\mathbf{x}} \sin \theta \cos \varphi+\overrightarrow{\mathbf{y}} \sin \theta \sin \varphi+\overrightarrow{\mathbf{z}} \cos \theta)$ |  |
| del | $\nabla$ | $\overrightarrow{\mathbf{x}} \frac{\partial}{\partial \mathrm{x}}+\overrightarrow{\mathbf{y}} \frac{\partial}{\partial \mathrm{y}}+\overrightarrow{\mathbf{z}} \frac{\partial}{\partial \mathrm{z}}$ |  |
| Laplacian (three dimensional, Cartesian coordinates) | $\nabla^{2}$ | $\frac{\partial^{2}}{\partial \mathrm{x}^{2}}+\frac{\partial^{2}}{\partial \mathrm{y}^{2}}+\frac{\partial^{2}}{\partial \mathrm{z}^{2}}$ |  |
| Laplacian (spherical coordinates) | $\nabla^{2}$ | $\frac{1}{r^{2}}\left(\frac{\partial}{\partial r} \mathrm{r}^{2} \frac{\partial}{\partial \mathrm{r}}\right.$ | $\left.\frac{1}{\sin \theta} \frac{\partial}{\partial \theta} \sin \theta \frac{\partial}{\partial \theta}+\frac{1}{\sin ^{2} \theta} \frac{\partial^{2}}{\partial \varphi^{2}}\right)$ |
| linear momentum (in x direction, for example) | $\hat{p}_{\mathrm{x}}$ | $-i \hbar \frac{\partial}{\partial \mathrm{x}}$ | There are also corresponding operators for $\mathrm{y}, \mathrm{z} .$ |


| linear momentum (3 dimensions) | $\hat{\mathbf{p}}$ | -i $\ddagger \nabla$ |  |
| :---: | :---: | :---: | :---: |
| $z$-component of angular momentum | $\hat{M}_{2}$ | $-i \hbar \frac{\partial}{\partial \varphi}$ |  |
| the square of total angular momentum | $\hat{M}^{2}$ | $-\hbar^{2}\left[\frac{1}{\sin \theta} \frac{\partial}{\partial \theta} \sin \theta \frac{\partial}{\partial \theta}+\frac{1}{\sin ^{2} \theta} \frac{\partial^{2}}{\partial \varphi^{2}}\right]$ |  |
| electronic angular momentum | $\hat{\mathbf{L}}$ | $\overrightarrow{\mathbf{x}} \hat{\mathrm{L}}_{\mathrm{x}}+\overrightarrow{\mathbf{y}} \hat{\mathrm{L}}_{\mathrm{y}}+\overrightarrow{\mathbf{z}} \hat{\mathrm{L}}_{\mathrm{z}}$ | Operator has the same form as the angular momentum operator, $\hat{\mathrm{M}}$ |
| total energy <br> (Hamiltonian) | $\hat{H}$ | $\hat{\mathrm{T}}+\hat{\mathrm{V}}$ | The potential energy (V) depends on the specific system being modeled. Additional energy terms can be added (e.g., interactions with electric or magnetic fields.) |
| kinetic energy for a single particle (Cartesian coordinates) | $\hat{\mathrm{T}}_{\mathrm{x}}$ | $\left(\frac{-\hbar^{2}}{2 m}\right) \frac{\partial^{2}}{\partial \mathrm{x}^{2}}$ | one-dimensional expression |
|  | $\hat{T}$ | $\left(\frac{-\hbar^{2}}{2 \mathrm{~m}}\right) \nabla^{2}$ | generalized expression |
| kinetic energy for a single particle (spherical coordinates) | $\hat{T}$ | $\frac{-\hbar^{2}}{2 \mu \mathrm{r}^{2}}\left[\frac{\partial}{\partial \mathrm{r}} \mathrm{r}^{2} \frac{\partial}{\partial \mathrm{r}}+\frac{1}{\sin \theta} \frac{\partial}{\partial \theta} \sin \theta \frac{\partial}{\partial \theta}+\frac{1}{\sin ^{2} \theta} \frac{\partial^{2}}{\partial \varphi^{2}}{ }_{-}\right.$ |  |


| potential energy for a particle in a box | $\hat{\mathrm{V}}$ | $\mathrm{V}=0$ <br> inside box $(0 \leq x \leq L)$ $V=\infty$ <br> outside box $(x<0, x>L)$ | The box can be extended to 2 or 3 dimensions and the length of the box, L , can be different in each direction. |
| :---: | :---: | :---: | :---: |
| harmonic oscillator potential | $\hat{\mathrm{V}}$ | $\frac{\mathrm{K}}{2}\left(\mathrm{x}-\mathrm{x}^{\mathrm{e}}\right)^{2}$ | For diatomic molecules; K is the force constant; $\mathrm{x}^{8}$ is the equilibrium bond length. |
|  |  | $\frac{K}{2} \hat{Q}^{2}$ | For polyatomic molecules; $\hat{Q}$ is the operator for the magnitude of the normal coordinate. |
| Morse (arharmonic) potential | $\hat{\mathrm{V}}$ | $D_{e}\left(1-e^{-\beta x}\right)^{2}$ | where x is the displacement of the oscillator from its equilibrium position, and $D_{e}$ and $\beta$ are constants or parameters. |
| Coulombic potential energy | $\hat{\mathrm{V}}$ | $\frac{q_{1} q_{2}}{4 \pi \varepsilon_{n} r_{12}}$ | q is the charge on each particle; r is the distance between particles. |
| electric field | $\hat{\mathbf{E}}$ | E | Simple multiplicative operator as long as the field is constant over the area of interaction with the system. |
| magnetic field | B | B | Operator is multiplicative and can be written without the caret. |


| magnetic moment of an electron | $\hat{H}_{\text {m }}$ | $-\frac{\mathrm{e}}{2 \mathrm{~m}_{\boldsymbol{e}}} \hat{\mathrm{L}}$ |  |
| :---: | :---: | :---: | :---: |
| energy of interaction with a magnetic field | $\hat{H}_{\text {II }}$ | $-\hat{\mu}_{\text {mil }} \bullet$ B | general form |
|  |  | $\frac{\mathrm{e}}{2 \mathrm{~m}} \mathrm{~L}_{\text {e }} \hat{\mathrm{L}} \bullet \mathrm{B}$ | the particle is an electron and the magnetic field is directed along the $z$-axis |
|  |  | $\frac{\mathrm{eB}}{2} \frac{\mathrm{~m}_{*}}{} \hat{\mathrm{~L}}_{\mathrm{z}}$ |  |
| spin orbit interaction | $\hat{H}_{s-0}$ | $\lambda \hat{\mathbf{S}} \cdot \hat{\mathrm{I}}$ |  |
| Coulomb operator | J | $\hat{J}_{j}(1) \varphi_{i}(1)=\left[\int \varphi_{i}^{*}(2) \frac{1}{r_{12}} \varphi_{i}(2) d \tau_{2}\right] \varphi_{i}(1)$ |  |
| Exchange operator | $\hat{\mathrm{K}}$ | $\hat{\mathrm{K}}_{\mathrm{j}}(1) \varphi_{\mathrm{i}}(1)=\left[\int \varphi_{\mathrm{j}}^{*}(2) \frac{1}{\mathrm{r}_{12}} \varphi_{\mathrm{i}}(2) d \tau_{2}\right] \varphi_{\mathrm{j}}(1)$ |  |
| Fock operator | $\hat{\mathrm{F}}$ | $\hat{H}^{0}+\sum_{j=1}^{N}\left(2 \hat{S}_{j}-\hat{\mathrm{K}}_{\mathrm{j}}\right)$ |  |
|  |  | $-\frac{\hbar^{2}}{2 m} \nabla^{2}-\frac{Z e^{2}}{4 \pi \varepsilon_{0} r}+\sum_{j-1}^{N}\left(2 \hat{\mathrm{~S}}_{j}-\hat{\mathrm{K}}_{\mathrm{j}}\right)$ |  |



## Time-dependent Schrödinger Equation

Without potential $E=T$
With potential $\mathrm{E}=\mathrm{T}+\mathrm{V}$

$$
\begin{aligned}
& \text { ih } \frac{\partial \Psi(\mathbf{r}, \mathbf{t})}{\partial \mathrm{t}}=-\frac{\mathbf{h}^{2}}{2 \mathrm{~m}} \Delta \Psi(\mathbf{r}, \mathbf{t}) \\
& \text { ih } \frac{\partial \Psi(\mathbf{r}, \mathbf{t})}{\partial \mathrm{t}}=-\frac{\mathbf{h}^{2}}{2 \mathrm{~m}} \Delta \Psi(\mathbf{r}, \mathbf{t})+V(\mathbf{r}, \mathbf{t}) \Psi(\mathbf{r}, \mathbf{t})
\end{aligned}
$$

Erwin Rudolf Josef Alexander Schrödinger
Austrian 1887-1961

## Schrödinger Equation for stationary states

$$
\begin{aligned}
& \text { iih } \frac{\partial \Psi(r, t)}{\partial t}=-\frac{h^{2}}{2 m} \Delta \Psi(\mathbf{r}, \mathbf{t})+V(\mathbf{r}, \mathbf{t}) \Psi(\mathbf{r}, \mathbf{t}) \\
& \Psi(\mathrm{r}, \mathrm{t})=\Psi(\mathrm{r}) \cdot \Psi(\mathrm{t})=\mathrm{A} \exp \left(\frac{-\mathrm{iEt}}{\mathrm{~h}}\right) \Psi(\mathrm{r})
\end{aligned}
$$

## Schrödinger Equation for stationary states

$$
-\frac{h^{2}}{2 \mathrm{~m}} \Delta \Psi(\mathrm{r})+\mathrm{V}(\mathrm{r}) \Psi(\mathrm{r})=\mathrm{E} \Psi(\mathrm{r})
$$

$\mathbb{H} \Psi(x, y, z)=E \Psi(x, y, z) \quad$ with $\quad \mathbb{H}=-\frac{h^{2}}{2 m} \Delta+\mathbb{V}$
$H$ is the hamiltonian


Half penny bridge in Dublin


Sir William Rowean Hamilton Irish 1805-1865

Chemistry is nothing but an application of Schrödinger Equation (Dirac)


$$
\langle\Psi \mid \Psi\rangle\langle\Psi| O|\Psi\rangle
$$

## Dirac notations

Paul Adrien Dirac 1902-1984
Dirac's mother was British and his father was Swiss65


## Uncertainty principle

the Heisenberg uncertainty principle states that locating a particle in a small region of space makes the momentum of the particle uncertain; and conversely, that measuring the momentum of a particle precisely makes the position uncertain

We already have seen incompatible operators
Werner Heisenberg
German
1901-1976

It is not surprising to find that quantum mechanics does not predict the position of an electron exactly. Rather, it provides only a probability as to where the electron will be found.
We shall illustrate the probability aspect in terms of the system of an electron confined to motion along a line of length $L$. Quantum mechanical probabilities are expressed in terms of a distribution function.
For a plane wave, p is defined and the position is not.
With a superposition of plane waves, we introduce an uncertainty on $p$ and we localize. Since, the sum of 2 wavefucntions is neither an eigenfunction for $p$ nor $x$, we have average values.
With a Gaussian function, the localization below is $1 / 2 \pi$

$p$ and $x$ do not commute and are incompatible For a plane wave, p is known and x is not ( $\Psi^{*} \Psi=\mathrm{A}^{2}$ everywhere) Let's superpose two waves... $\lambda_{1}=\lambda+\frac{\Delta \lambda}{2} \operatorname{et} \lambda_{2}=\lambda-\frac{\Delta \lambda}{2}$
this introduces a delocalization tor $p$ and may be localize $x$
At the origin $\mathrm{x}=0$ and at $\mathrm{t}=0$ we want to increase the total amplitude, so the two waves $\Psi_{1}$ and $\Psi_{2}$ are taken in phase At $\pm \Delta x / 2$ we want to impose them out of phase The position is therefore known for $\mathrm{x} \pm \Delta \mathrm{x} / 2$ the waves will have wavelengths

$$
\begin{aligned}
& \frac{\mathrm{x}}{\lambda_{1}}-\frac{\mathrm{x}}{\lambda_{2}}=\frac{1}{2} \quad \rightarrow \quad \frac{\frac{\Delta \mathrm{x}}{2}}{\lambda+\frac{\Delta \lambda}{2}} \cdot \frac{\frac{\Delta \mathrm{x}}{2}}{\lambda \cdot \frac{\Delta \lambda}{2}}=\frac{1}{2} \\
& \frac{\frac{\Delta \mathrm{x}}{2}\left(\lambda-\frac{\Delta \lambda}{2}\right) \cdot \frac{\Delta \mathrm{x}}{2}\left(\lambda+\frac{\Delta \lambda}{2}\right)}{\left(\lambda+\frac{\Delta \lambda}{2}\right)\left(\lambda \cdot \frac{\Delta \lambda}{2}\right)}=\frac{1}{2}=\frac{-\Delta \mathrm{x} \Delta \lambda}{2 \lambda^{2}} \\
& \Delta \mathrm{x} \Delta \lambda=-\lambda^{2} \quad \mathrm{p}=\frac{\mathrm{h}}{\lambda} \rightarrow \mathrm{dp}=\frac{-\mathrm{hd} \lambda}{\lambda^{2}} \\
& \underset{\rightarrow}{\text { making } \Delta \mathrm{p}=\mathrm{dp} \text { and } \Delta \lambda=\mathrm{d} \lambda} \\
& \Delta \mathrm{x} \cdot \Delta \mathrm{p}=\mathrm{h}
\end{aligned}
$$

## Superposition of two waves

$$
\begin{aligned}
& \Psi=\cos \left(2 \pi\left(\frac{\mathrm{x}}{0.9}-v \mathrm{t}\right)\right)+\cos \left(2 \pi\left(\frac{\mathrm{x}}{1.1}-v \mathrm{t}\right)\right)=2 \cos \left(\frac{\pi \mathrm{x}}{9.9}\right) \cos \left(2 \pi\left(\frac{\mathrm{x}}{0.99}-v \mathrm{t}\right)\right) \\
& \frac{\Delta \mathrm{x}}{2}=4.95 \quad \Delta \lambda=1.1-0.9=0.2 \\
& \Delta x \cdot \Delta \lambda=0.99 \text { close to } \lambda^{2}=1
\end{aligned}
$$



Factor $1 / 2 \pi$ a more realistic localization


## Uncertainty principle

A more accurate calculation localizes more ( $1 / 2 \pi$ the width of a gaussian) therefore one gets


Werner Heisenberg

German
1901-1976
$x$ and $p$ or $E$ and $t$ play symmetric roles
in the plane wave expression;
Therefore, there are two main uncertainty principles

